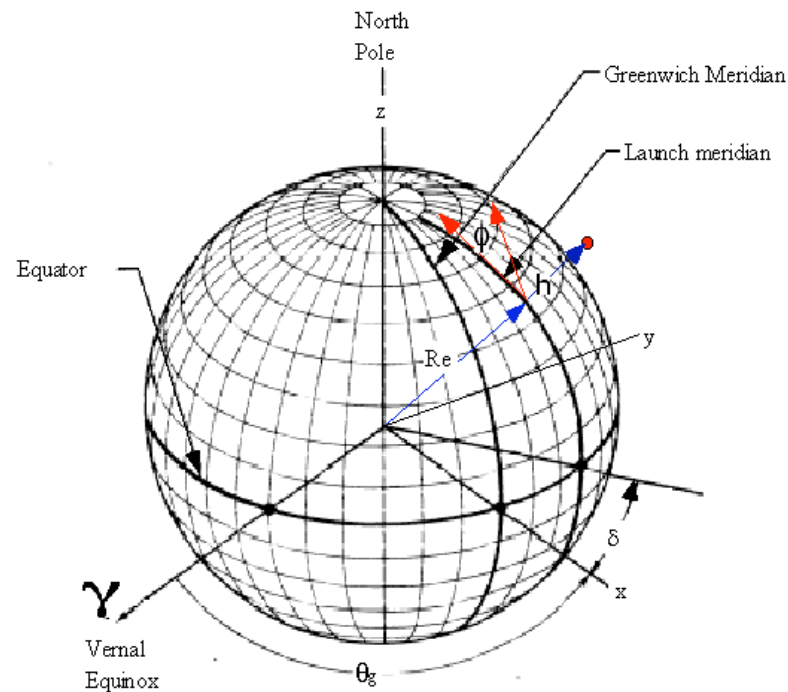


Appendix to Section 3: A brief overview of Geodetics

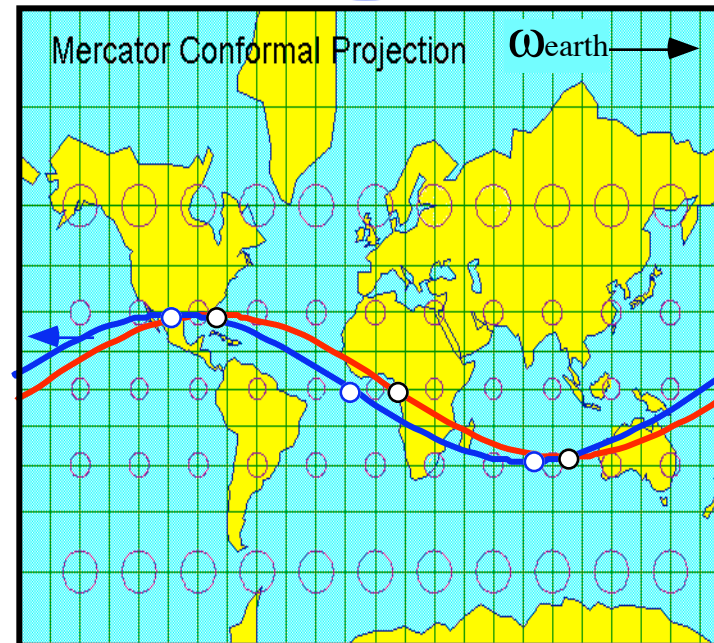
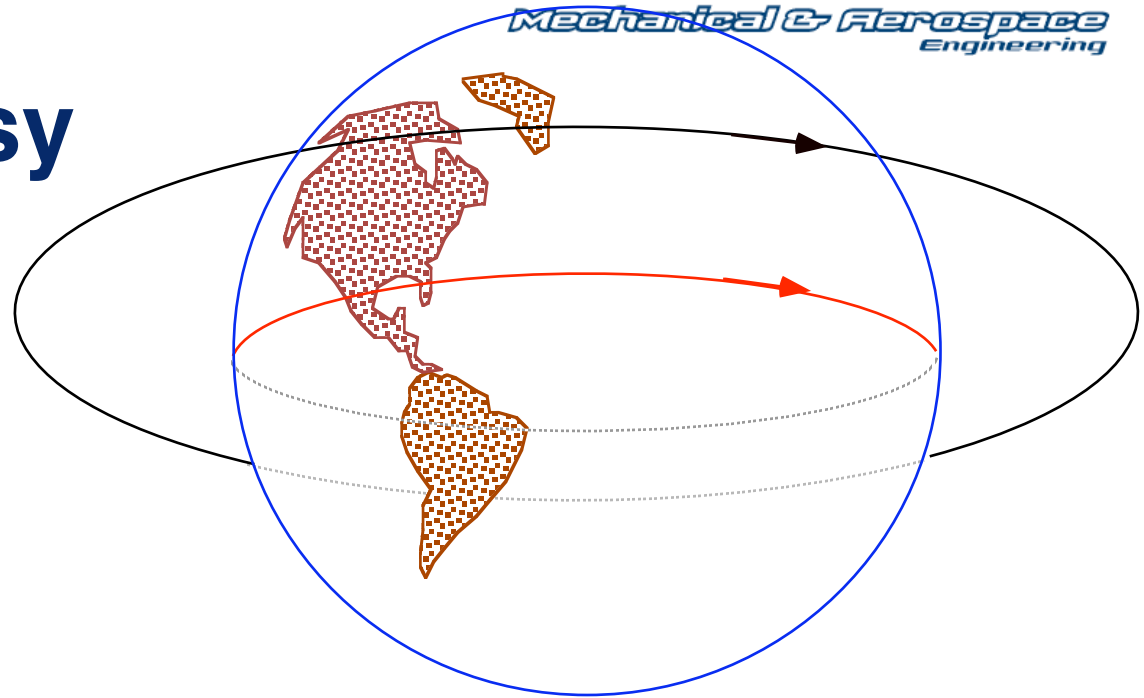


Geodesy

- Navigation Geeks do Calculations in Geocentric (spherical) Coordinates
- Map Makers Give Surface Data in Terms of Geodetic (elliptical) Coordinates
- Need to have some idea how to relate one to another

-- *science of geodesy*

MAE 5540 - Propulsion Systems



How Does the Earth Radius Vary with Latitude?

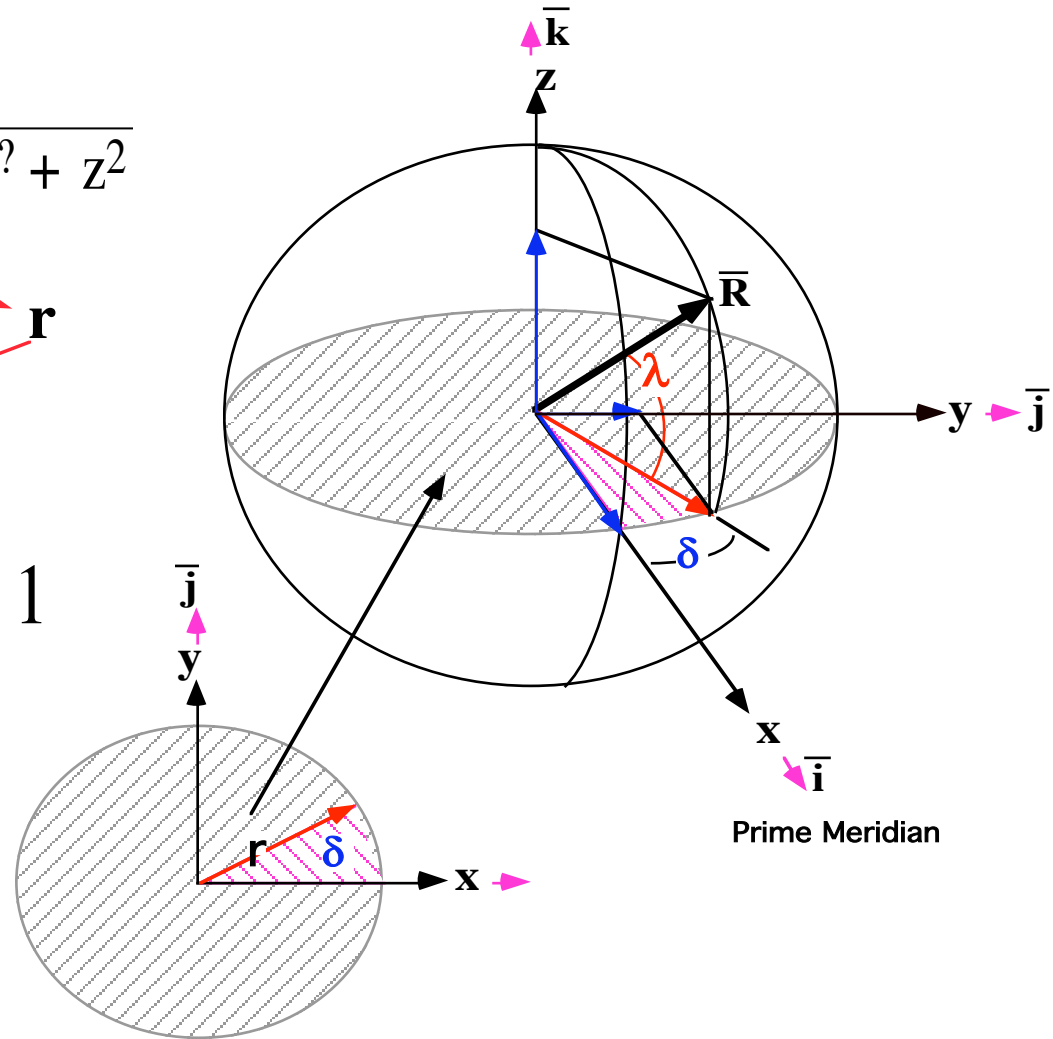
$$R_{\text{Earth}} = \sqrt{x^2 + y^2 + z^2} = \sqrt{r^2 + z^2}$$

Ellipse:

$$\left[\frac{r}{R_{\text{eq}}} \right]^2 + \left[\frac{z}{R_{\text{eq}} \sqrt{1 - e_{\text{Earth}}^2}} \right]^2 = 1$$

"a"

"b"



How Does the Earth Radius vary with Latitude?

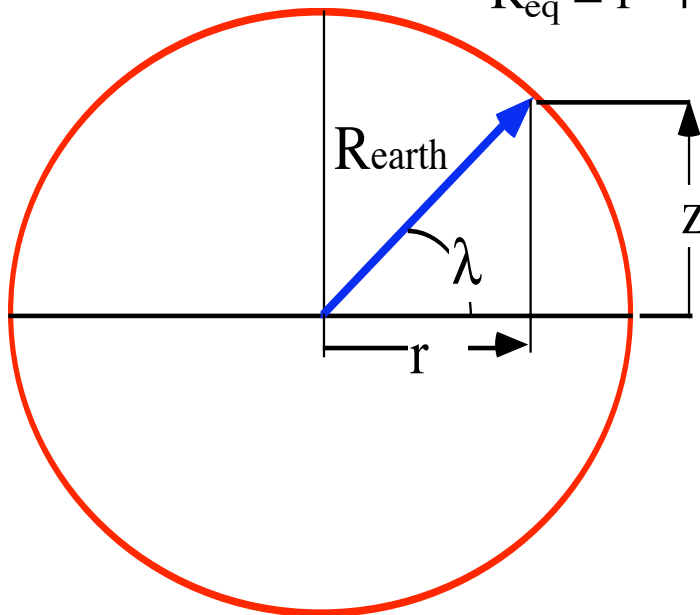
$$\left[\frac{r}{R_{eq}} \right]^2 + \left[\frac{z}{R_{eq} \sqrt{1 - e_{Earth}^2}} \right]^2 = 1$$

⇓

$$[1 - e_{Earth}^2] r^2 + z^2 = R_{eq}^2 [1 - e_{Earth}^2]$$

⇓

$$R_{eq}^2 = r^2 + \frac{z^2}{[1 - e_{Earth}^2]} = r^2 \left[1 + \frac{\left[\frac{z}{r} \right]^2}{[1 - e_{Earth}^2]} \right]$$



$$r^2 = R_{earth}^2 \cos^2[\lambda]$$

$$\left[\frac{z}{r} \right]^2 = \tan^2[\lambda]$$

How Does the Earth Radius vary with Latitude?

$$\frac{R_{eq}^2}{R_{Earth}^2} = \cos^2 [\lambda] \left[1 + \frac{\tan^2 [\lambda]}{[1 - e_{Earth}^2]} \right] =$$

$$\left[\frac{[1 - e_{Earth}^2] \cos^2 [\lambda] + \sin^2 [\lambda]}{[1 - e_{Earth}^2]} \right] =$$

$$\frac{\cos^2 [\lambda] + \sin^2 [\lambda] - e_{Earth}^2 \cos^2 [\lambda]}{[1 - e_{Earth}^2]} = \frac{1 - e_{Earth}^2 \cos^2 [\lambda]}{[1 - e_{Earth}^2]}$$

Inverting

$$\frac{R_{earth}(\lambda)}{R_{eq}} = \sqrt{\frac{1 - e_{Earth}^2}{1 - e_{Earth}^2 \cos^2 [\lambda]}}$$

Earth Radius vs Geocentric Latitude

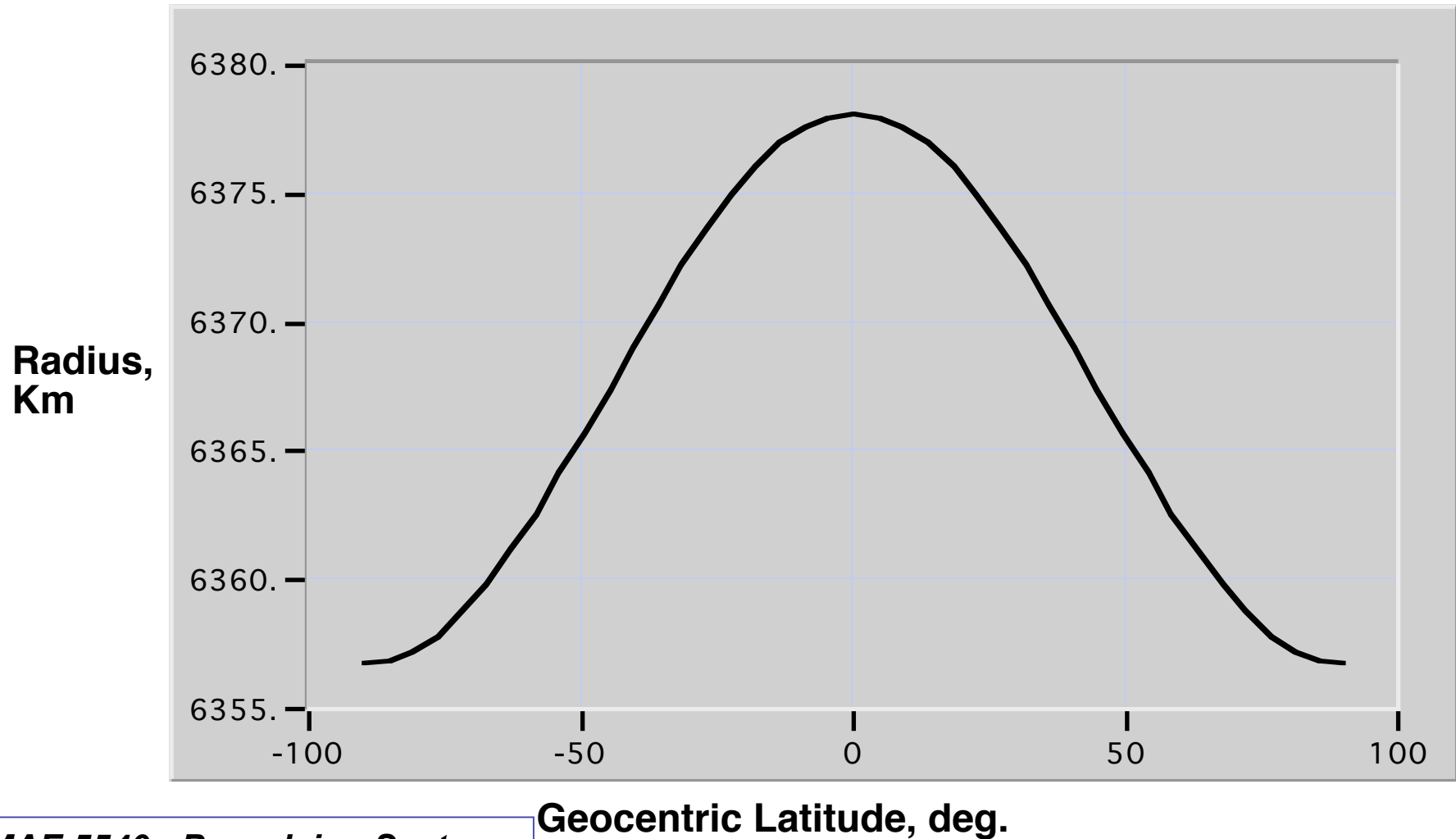
$$\frac{R_{\text{earth}(\lambda)}}{R_{\text{eq}}} = \sqrt{\frac{1 - e_{\text{Earth}}^2}{1 - e_{\text{Earth}}^2 \cos^2 [\lambda]}}$$

Polar Radius: 6356.75170 km
Equatorial Radius: 6378.13649 km

$$e_{\text{Earth}} = \sqrt{1 - \left[\frac{b}{a}\right]^2} = \sqrt{\frac{a^2 - b^2}{a^2}} =$$

$$\frac{\sqrt{[6378.13649]^2 - 6378.13649^2}}{[6378.13649]} = 0.08181939$$

Earth Radius vs Geocentric Latitude (concluded)



Earth Radius ... alternate formula

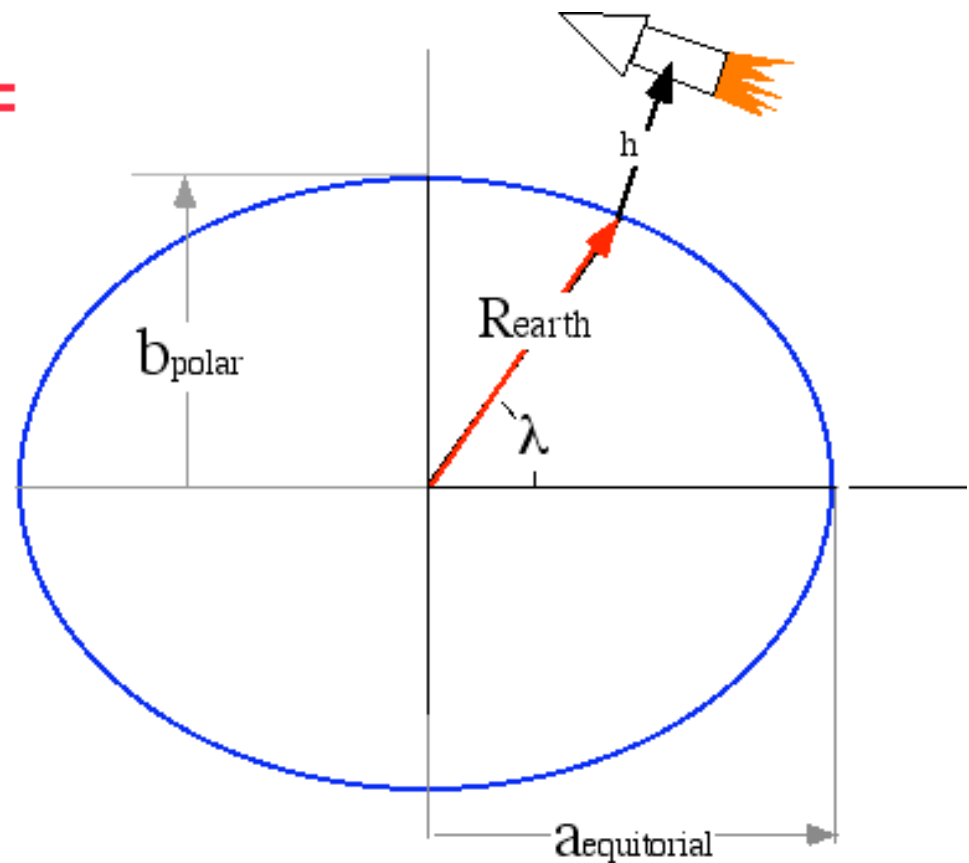
- Earth radius as Function of Latitude

$$R_{\text{earth}} = \frac{a_{\text{equitorial}}}{\sqrt{1 + \frac{e_{\text{earth}}^2}{1 - e_{\text{earth}}^2} \sin^2 \lambda}}$$

$$a_{\text{equitorial}} = 6378.13649 \text{ km}$$

$$b_{\text{polar}} = 6356.7515 \text{ km}$$

$$e_{\text{earth}} = \sqrt{1 - \left[\frac{b_{\text{polar}}}{a_{\text{equitorial}}} \right]^2}$$



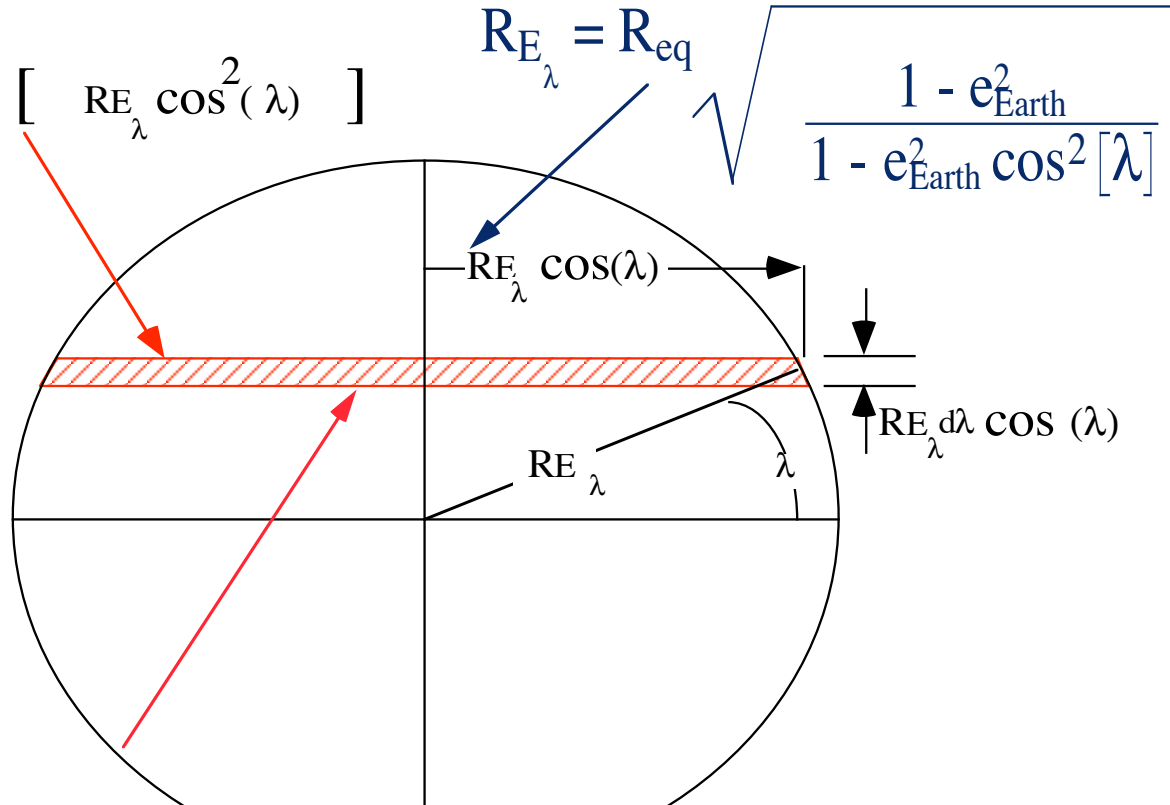
What is the mean radius of the earth?

$$dA = \pi \left[R_{E\lambda} \cos^2(\lambda) \right]$$

$$R_{E\lambda} = R_{eq} \sqrt{\frac{1 - e_{Earth}^2}{1 - e_{Earth}^2 \cos^2[\lambda]}}$$

**IAU Convention:
Based on
Earth's Volume**

Sphere Volume:

$$\frac{4\pi}{3} R_{E_{mean}}^3 = V_E$$


$$dV = \pi \left[R_{E\lambda} \cos^2(\lambda) \right] \times R_{E\lambda} d\lambda \cos(\lambda)$$

What is the Earth's Mean Radius?

(continued)

- **Earth's (ellipsoid) Volume**

$$V_E = \int_{-\frac{\pi}{2}}^{\frac{\pi}{2}} \pi \left[R_{E\lambda} \cos(\lambda) \right]^3 d\lambda =$$

$$R_{eq}^3 \int_{-\frac{\pi}{2}}^{\frac{\pi}{2}} \pi \left[\frac{1 - e_{Earth}^2}{1 - e_{Earth}^2 \cos^2[\lambda]} \right]^{3/2} \cos^3(\lambda) d\lambda =$$

$$\frac{4\pi}{3} \sqrt{1 - e_{earth}^2} R_{eq}^3$$

What is the Earth's Mean Radius?

(continued)

- **Based on Volume**

$$\text{Ellipsoid Volume: } \frac{4\pi}{3} \sqrt{1-e^2} R_{\text{eq}}^3$$

$$\text{Sphere Volume: } \frac{4\pi}{3} R_{\text{sphere}}^3$$



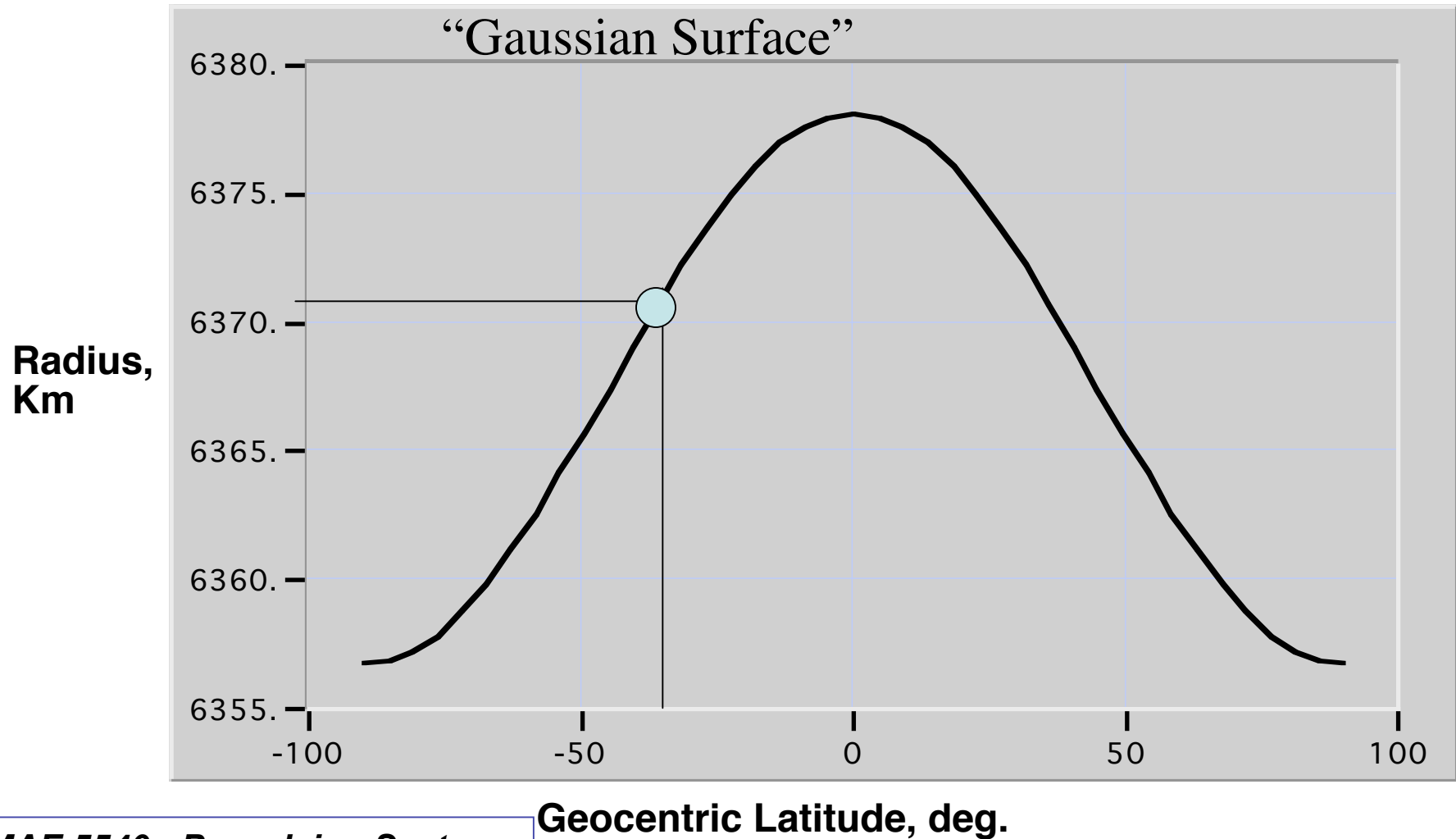
$$R_{\text{sphere}} \approx R_{\text{mean}} = [1 - 0.08181939^2]^{1/6} 6378.13649 = 6371.0002 \text{ km}$$

- **Mean Radius** we have been using is for a Sphere with same volume as the Earth

$$M_E = \rho_E V_E$$

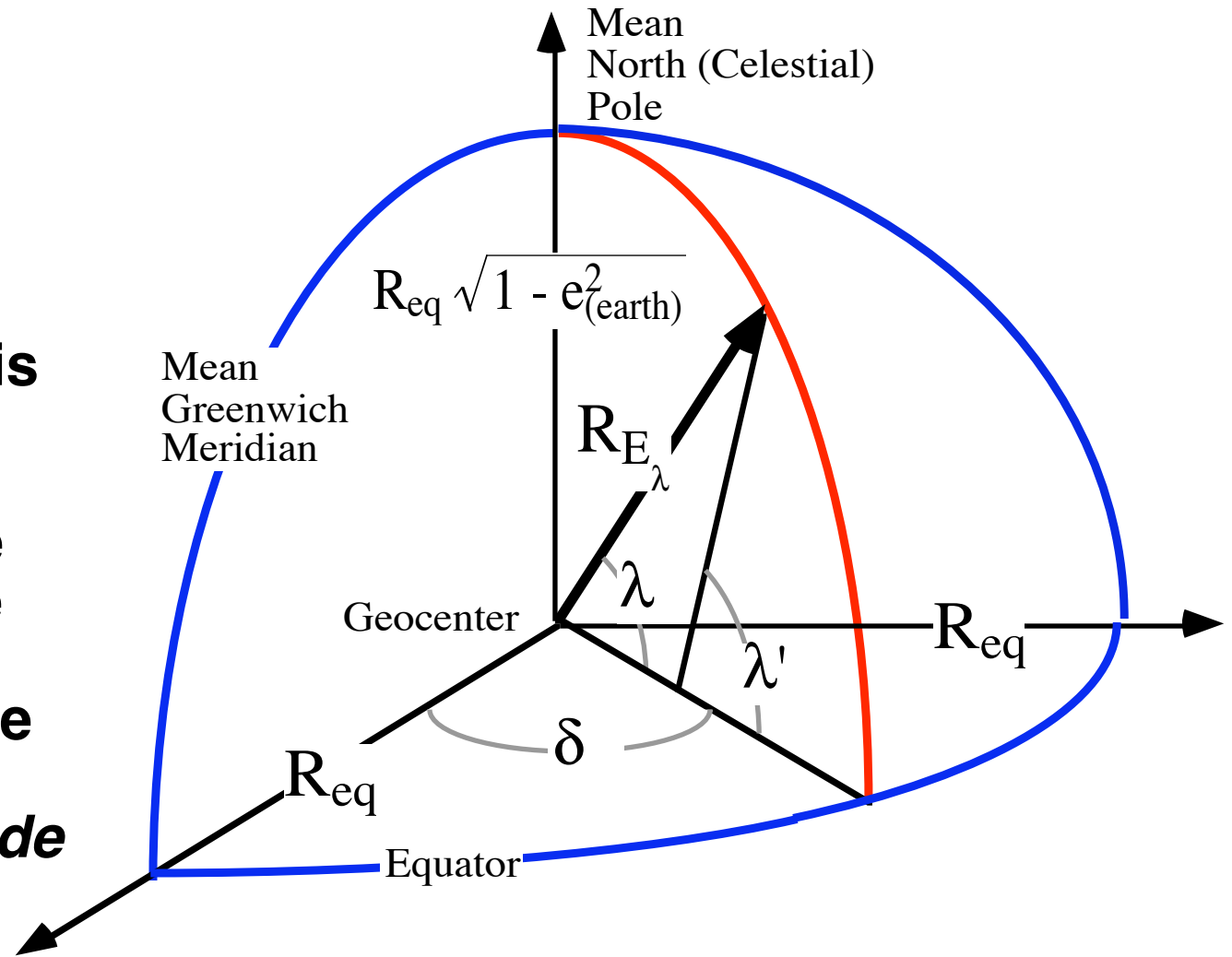
"gravitational radius"

Earth Radius vs Geocentric Latitude (concluded)



Geocentric vs Geodetic Coordinates

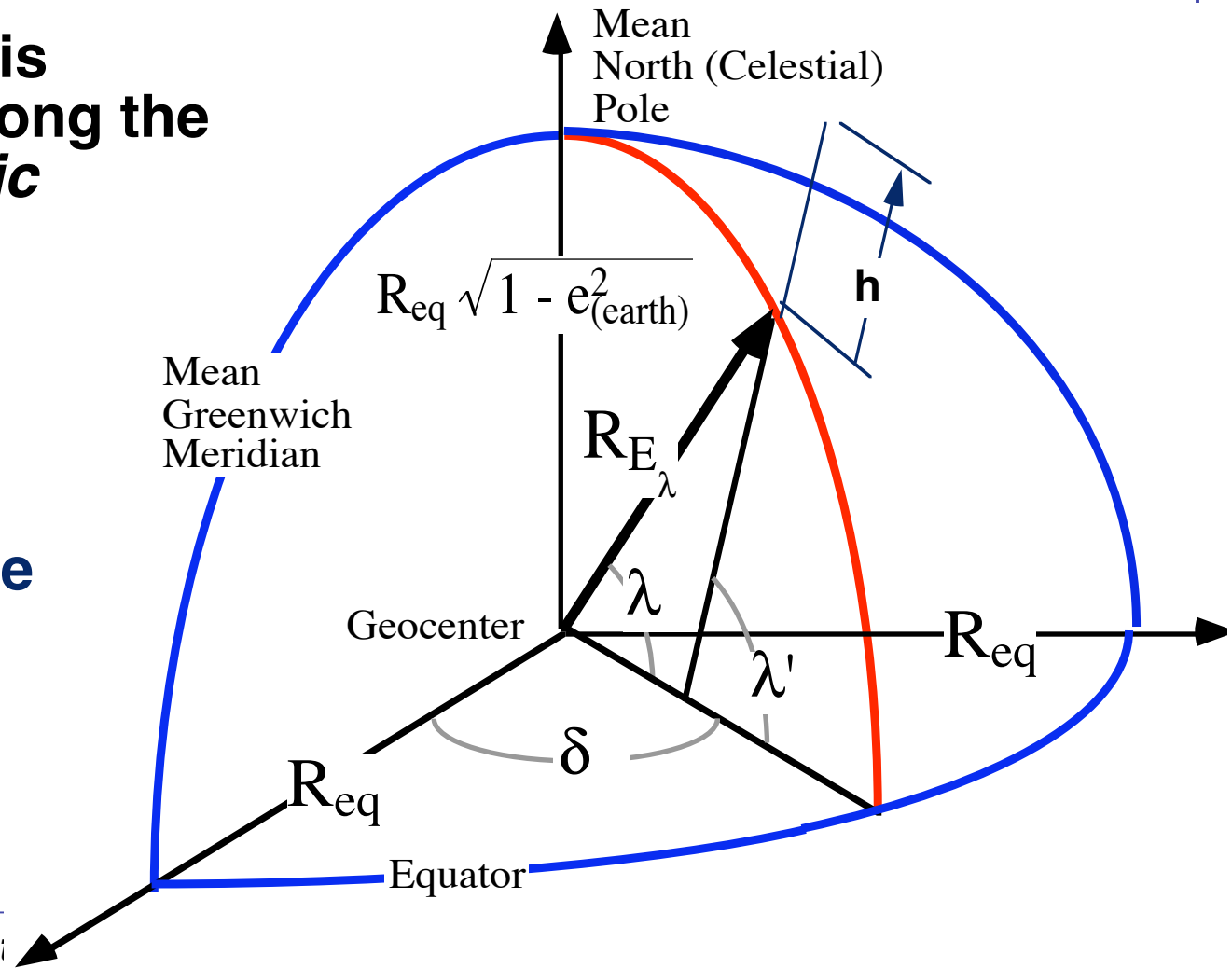
- Map makers define a new latitude which is the angle that normal to the Earth's surface makes with the equatorial plane
- *Geodetic latitude*



Geocentric vs Geodetic Coordinates *(continued)*

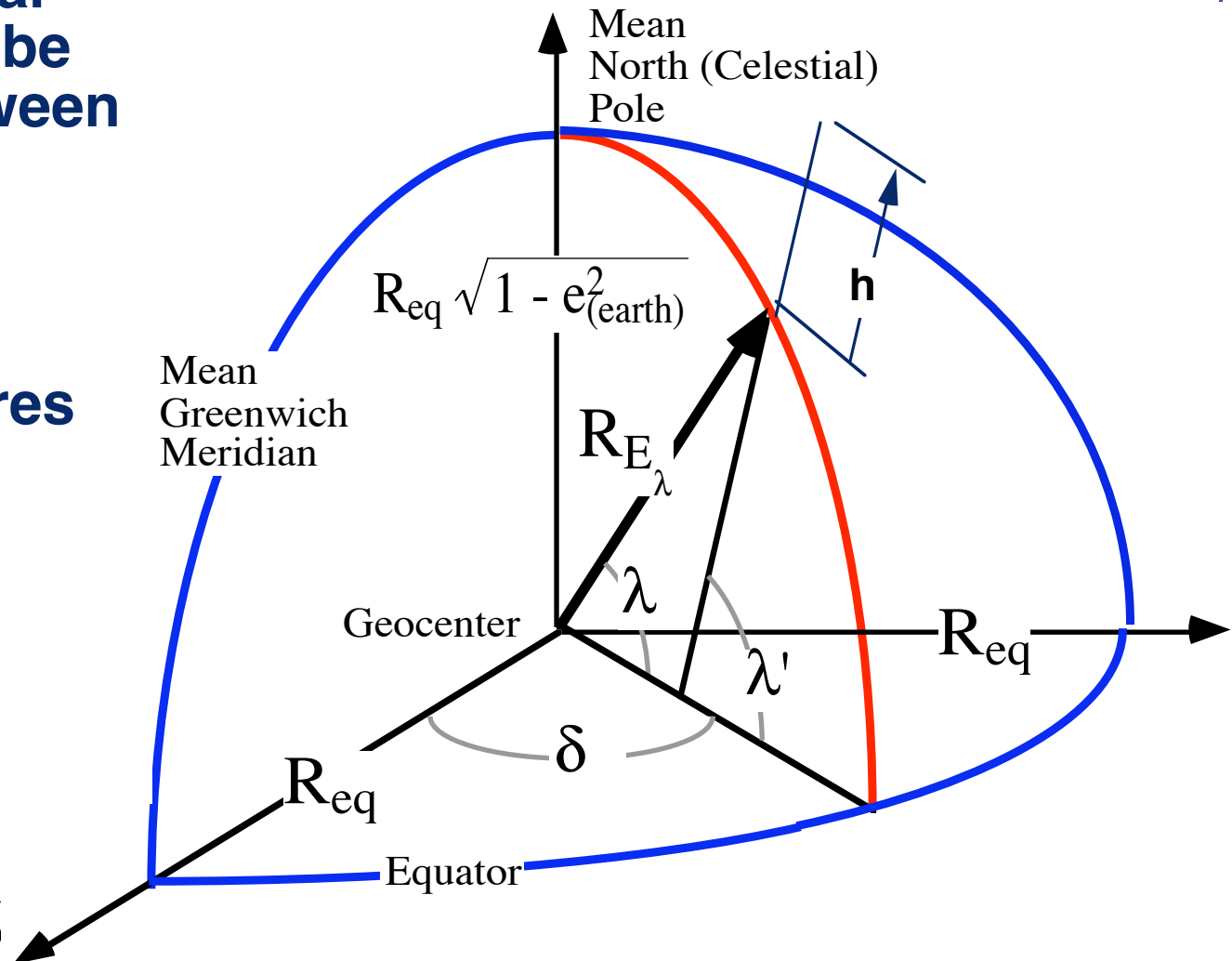
- Since the Earth is Elliptical only along the z-axis ... *geodetic* and *geocentric longitude* are identical

- Altitude is an extension of the *line of latitude geodetic*



Geocentric vs Geodetic Coordinates *(continued)*

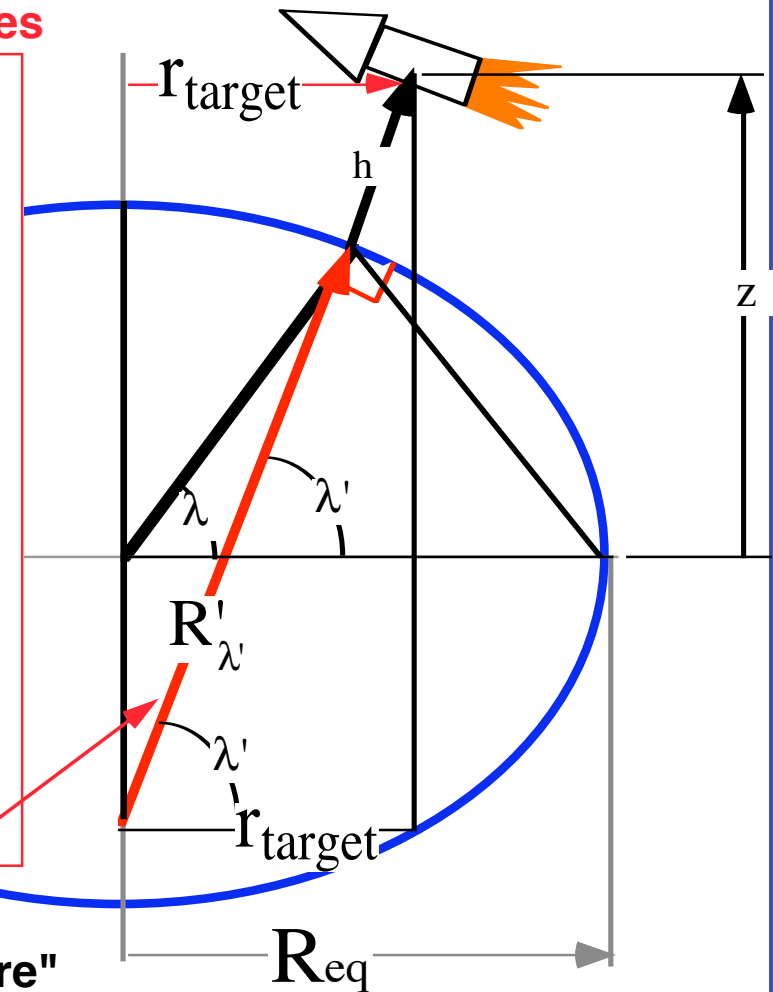
- Complex nonlinear equations describe relationship between geocentric and geodetic latitude
- Derivation requires Extensive Knowledge of Spherical Trigonometry



Geocentric vs Geodetic Coordinates *(continued)*

Geocentric Cartesian Coordinates

$$\begin{aligned}
 * \quad x_{\text{target}} &= [R'_{\lambda'} + h] \cos(\lambda') \cos(\delta) \\
 y_{\text{target}} &= [R'_{\lambda'} + h] \cos(\lambda') \sin(\delta) \\
 z_{\text{target}} &= [R'_{\lambda'} [1 - e^2_{\text{earth}}] + h] \sin(\lambda') \\
 R'_{\lambda'} &= \frac{R_{\text{eq}}}{\sqrt{1 - e^2_{\text{earth}} \sin^2(\lambda')}}
 \end{aligned}$$



* We would be here all week
if I try to derive this

"Radius of Curvature"

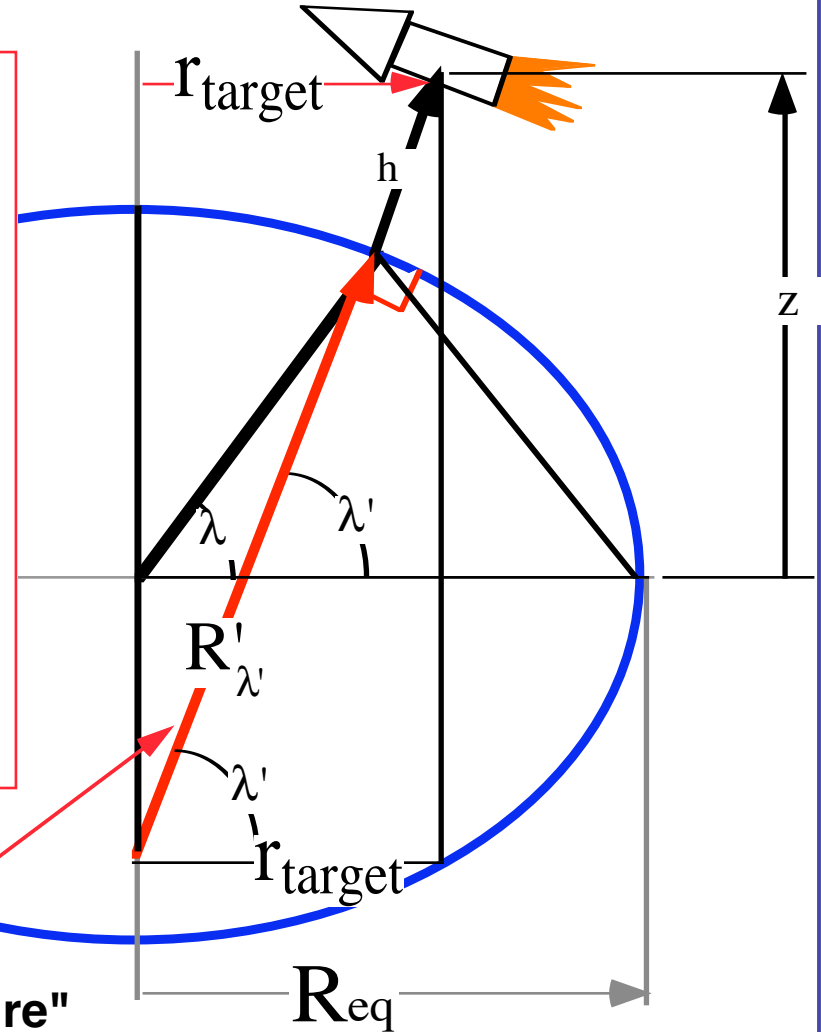
Geocentric vs Geodetic Coordinates *(continued)*

Geocentric Polar Coordinates

$$R_{\text{target}} = \sqrt{x_{\text{target}}^2 + y_{\text{target}}^2 + z_{\text{target}}^2}$$

$$\delta_{\text{target}} = \tan^{-1} \left[\frac{y_{\text{target}}}{x_{\text{target}}} \right]$$

$$\lambda_{\text{target}} = \tan^{-1} \left[\frac{z_{\text{target}}}{\sqrt{x_{\text{target}}^2 + y_{\text{target}}^2}} \right]$$



"Radius of Curvature"

Geocentric vs Geodetic Coordinates *(concluded)*

Inverse Relationships, non-linear no direct solution

$$h = \frac{\sqrt{X_{\text{target}}^2 + Y_{\text{target}}^2}}{\cos(\lambda')} - R'_{\lambda'}$$

$$\delta_{\text{target}} = \tan^{-1} \left[\frac{Y_{\text{target}}}{X_{\text{target}}} \right]$$

$$\lambda'_{\text{target}} = \tan^{-1} \left[\frac{Z_{\text{target}}}{\sqrt{X_{\text{target}}^2 + Y_{\text{target}}^2}} \times \left[\frac{1}{1 - e_{\text{earth}}^2 \frac{R'_{\lambda'}}{R'_{\lambda'} + h_{\text{target}}}} \right] \right]$$

Whitmore, Stephen A., and Haering, Edward A., Jr., FORTRAN Program for the Analysis of Ground Based Range Tracking Data--Usage and Derivations, NASA TM 104201, December, 1992

Pulling it all together

- **Given geodetic coordinates -- compute geocentric**

i) Compute geocentric cartesian coordinates

Range, runway thresholds, radar antennae, beacon

$$X_{\text{target}} = [R'_{\lambda'} + h] \cos(\lambda') \cos(\delta)$$

$$Y_{\text{target}} = [R'_{\lambda'} + h] \cos(\lambda') \sin(\delta)$$

$$Z_{\text{target}} = [R'_{\lambda'} [1 - e^2_{\text{earth}}] + h] \sin(\lambda')$$

$$R'_{\lambda'} = \frac{R_{\text{eq}}}{\sqrt{1 - e^2_{\text{earth}} \sin^2(\lambda')}}}$$

Pulling it all together (continued)

- **Given geodetic coordinates -- compute geocentric**
 - ii) Compute Geocentric polar coordinates next

$$R_{\text{target}} = \sqrt{x_{\text{target}}^2 + y_{\text{target}}^2 + z_{\text{target}}^2}$$

$$\delta_{\text{target}} = \tan^{-1} \left[\frac{y_{\text{target}}}{x_{\text{target}}} \right]$$

$$\lambda_{\text{target}} = \tan^{-1} \left[\frac{z_{\text{target}}}{\sqrt{x_{\text{target}}^2 + y_{\text{target}}^2}} \right]$$

Pulling it all together (concluded)

- Given geocentric (usually x,y,z) coordinates -- compute geodetic

GPS, INS, TLE's

**No explicit solution:
requires**

- 1) series expansion solution,
- 2) numerical iteration,
- 3) or a special solution called "Ferrari's method"***

$$R'_{\lambda'} = \frac{R_{eq}}{\sqrt{1 - e_{earth}^2 \sin^2(\lambda')}}$$

$$h = \frac{\sqrt{X_{target}^2 + Y_{target}^2}}{\cos(\lambda')} - R'_{\lambda'}$$

$$\delta_{target} = \tan^{-1} \left[\frac{Y_{target}}{X_{target}} \right]$$

$$\lambda'_{target} = \tan^{-1} \left[\frac{Z_{target}}{\sqrt{X_{target}^2 + Y_{target}^2}} \times \left[\frac{1}{1 - e_{earth}^2 \frac{R'_{\lambda'}}{R'_{\lambda'} + h_{target}}} \right] \right]$$

***NASA Technical Paper 3430, Whitmore and Haering, FORTRAN Program for Analyzing Ground-Based Tracking Data: Usage and Derivations, Version 6.2, 1995

Numerical Example

- **Edwards Air Force Base, Radar Site #34**

$$\lambda' = 34.96081^\circ$$

$$\delta = -117.91150^\circ$$

$$h = 2563.200 \text{ ft}$$

- **Find corresponding geocentric cartesian and polar coordinates**

Numerical Example (cont'd)

- **Compute Local Radius of Curvature**

$$R'_{\lambda'} = \frac{R_{eq}}{\sqrt{1 - e_{earth}^2 \sin^2(\lambda')}} =$$

$$\frac{6378.13649 \text{ km}}{\sqrt{1 - \left[0.08181939 \sin\left(34.96081 \times \frac{\pi}{180}\right)\right]^2}} =$$

$$6392.187109 \text{ km}$$

Numerical Example (cont'd)

- Compute X and Y (geocentric)

$$r_{\text{target}} = [R'_{\lambda'} + h] \cos(\lambda') =$$

$$[6392.1871 + (2536.2 \times 3.048 \times 10^{-4})] \cos\left(34.96081 \times \frac{\pi}{180}\right) =$$

$$5239.3131 \text{ km}$$

$$X_{\text{target}} = r_{\text{target}} \cos(\delta) =$$

$$5239.3131 \text{ km} \times \cos\left(-117.91150 \times \frac{\pi}{180}\right) =$$

$$-2452.5602 \text{ km}$$

$$Y_{\text{target}} = r_{\text{target}} \sin(\delta) =$$

$$5239.3131 \text{ km} \times \sin\left(-117.91150 \times \frac{\pi}{180}\right) =$$

$$-4629.83218 \text{ km}$$

Numerical Example (cont'd)

- Compute z (geocentric)

$$z_{\text{target}} = \left[R'_{\lambda'} \left[1 - e_{\text{earth}}^2 \right] + h \right] \sin(\lambda') =$$

$$\left[6392.1871 \text{ km} \left[1 - 0.08181939^2 \right] + (2536.2 \times 3.048 \times 10^{-4}) \right] \sin \left(34.96081 \times \frac{\pi}{180} \right) =$$

$$3638.7480 \text{ km}$$

- **Compute Geocentric Polar Coordinates**

$$R_{\text{target}} = \sqrt{2452.5602^2 + 4629.83218^2 + 3638.7480^2} = 6378.94104\text{km}$$

$$\delta_{\text{target}} = \tan^{-1} \left[\frac{y_{\text{target}}}{x_{\text{target}}} \right] =$$

$$\frac{180}{\pi} \times \tan^{-1} \left[\frac{-4629.83218}{-2452.5602} \right] = -117.9115^\circ$$

$$\lambda_{\text{target}} = \tan^{-1} \left[\frac{z_{\text{target}}}{\sqrt{x_{\text{target}}^2 + y_{\text{target}}^2}} \right] =$$

$$\frac{180}{\pi} \times \tan^{-1} \left[\frac{3638.7480}{5239.3131} \right] = 34.7803^\circ$$

Numerical Example (cont'd)

- Compute Local Earth Radius and Geocentric Distance Above Geoid

$$R_{E\lambda} = R_{eq} \sqrt{\frac{1 - e_{Earth}^2}{1 - e_{Earth}^2 \cos^2[\lambda]}} =$$

$$6378.13649 \text{ km} \sqrt{\frac{1 - 0.08181939^2}{1 - 0.08181939^2 \cos^2\left[34.7083 \times \frac{\pi}{180}\right]}} =$$

$$6364.23 \text{ km}$$

$$D_{geoid} = R_{target} - R_{E\lambda} =$$

$$[6378.94104 - 6364.23] \text{ km} = 14.7 \text{ km} = 48228.25 \text{ ft}$$

Numerical Example (concluded)

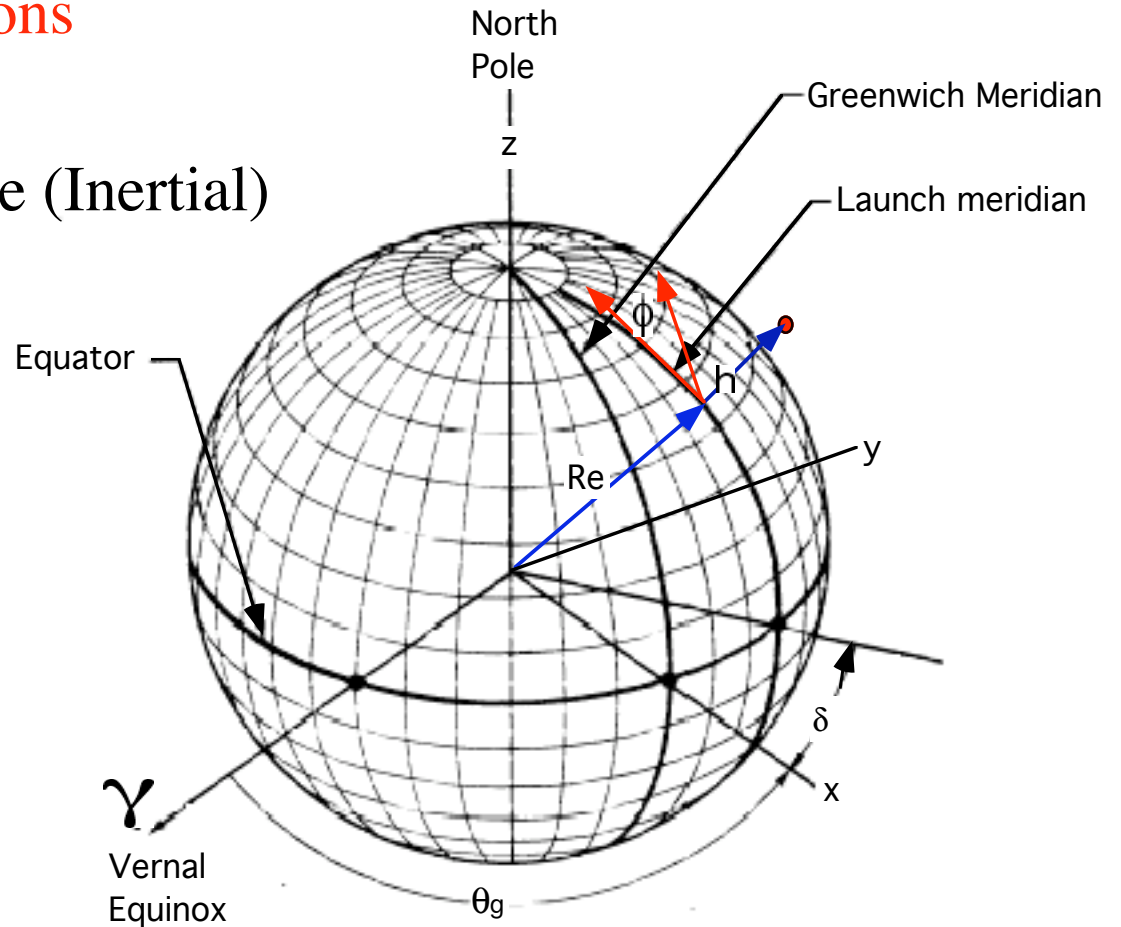
- Comparison

Geodetic		Geocentric
$\lambda' = 34.96081^\circ$	↔	$\lambda = 34.7803^\circ$
$\delta = -117.91150^\circ$		$\delta = -117.91150^\circ$
$h = 2563.200 \text{ ft}$	↔	$D_{\text{geoid}} = 48228.25 \text{ ft}$

- Earth Oblateness is NOT trivial, and in the REAL World -- it *must be* accounted for

Appendix II: Rigorous Derivation of Realizable Launch Inclination

- **Launch Initial Conditions**
- Position: λ , Latitude
 Ω , Longitude (Inertial)
 h , Altitude

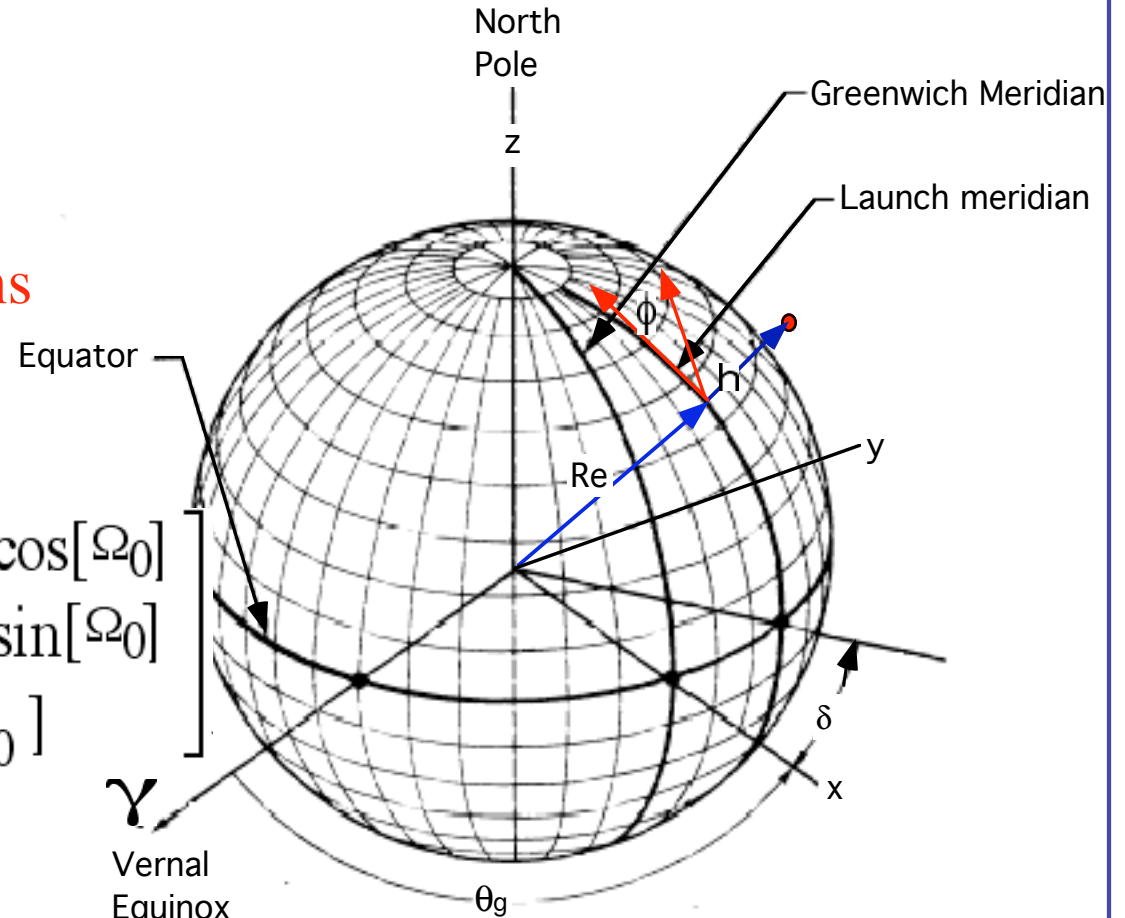


• Launch Initial Conditions

• Launch Initial Conditions
(Inertial Coordinates)

$$\begin{bmatrix} x_I \\ y_I \\ z_I \end{bmatrix}_0 = \begin{bmatrix} [R_{\text{earth}} + h_0] \cos[\lambda_0] \cos[\Omega_0] \\ [R_{\text{earth}} + h_0] \cos[\lambda_0] \sin[\Omega_0] \\ [R_{\text{earth}} + h_0] \sin[\lambda_0] \end{bmatrix}$$

$$\Omega_0 = \delta_0 + \theta_g$$



Sidereal hour angle

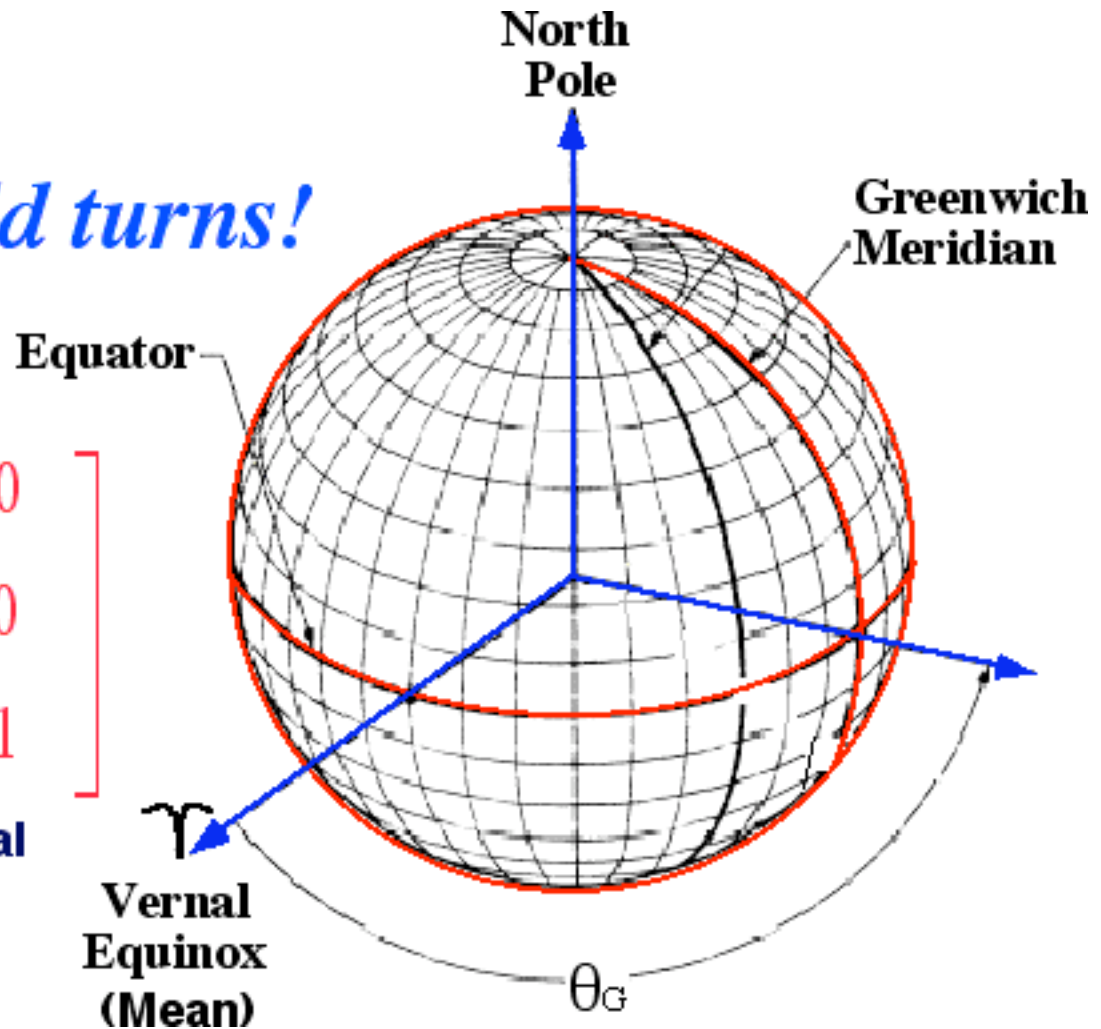
As the world turns!

$M_{\text{Rotation}} =$

$$\begin{bmatrix} \cos[\theta_G] & -\sin[\theta_G] & 0 \\ \sin[\theta_G] & \cos[\theta_G] & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

**Greenwich Sidereal
(Hour) Angle**

Ignore for now!



Computing the Hour Angle

- θ_G Historically Expressed in Hours

... Sometimes referred to as Greenwich Mean Sidereal Time ... but we are going to treat it as an angle

$$\theta_G = \omega_{\text{earth}} \times [T_{\text{GMST}} - T_{\text{JD2000}}]$$


- Sidereal time is a measure of the Earth's rotation with respect to distant celestial objects.